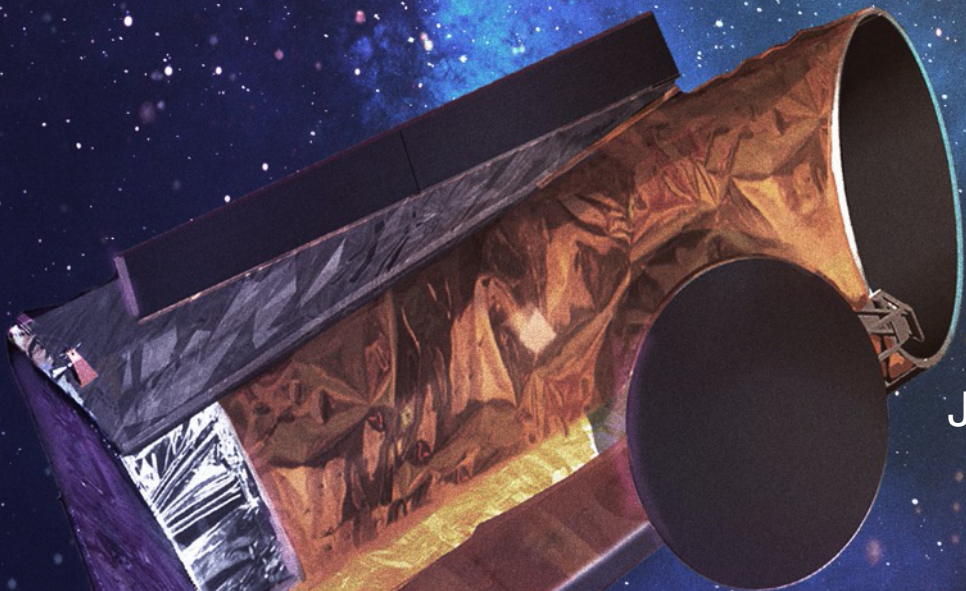


HabEx



The Habitable Exoplanet Observatory Concept

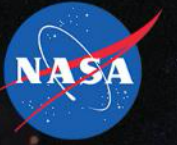
Exploring planetary systems around our neighboring sunlike stars and enabling a broad range of observatory science in the UV through the near-IR



Alina Kiessling

Jet Propulsion Laboratory / California Institute of Technology
Deputy Center Study Scientist

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- Through a community-driven competed and funded Guest Observer program, HabEx will enable extensive physics of the cosmos, cosmic origins, exoplanet exploration, and solar system science.
- To be sensitive to an uncertain future landscape, a matrix of 9 architectures of varying scope have been investigated.
- The HabEx concept is an exciting design that is budget conscious, low risk, and technologically implementable in the 2020s.



HabEx



Science Goals



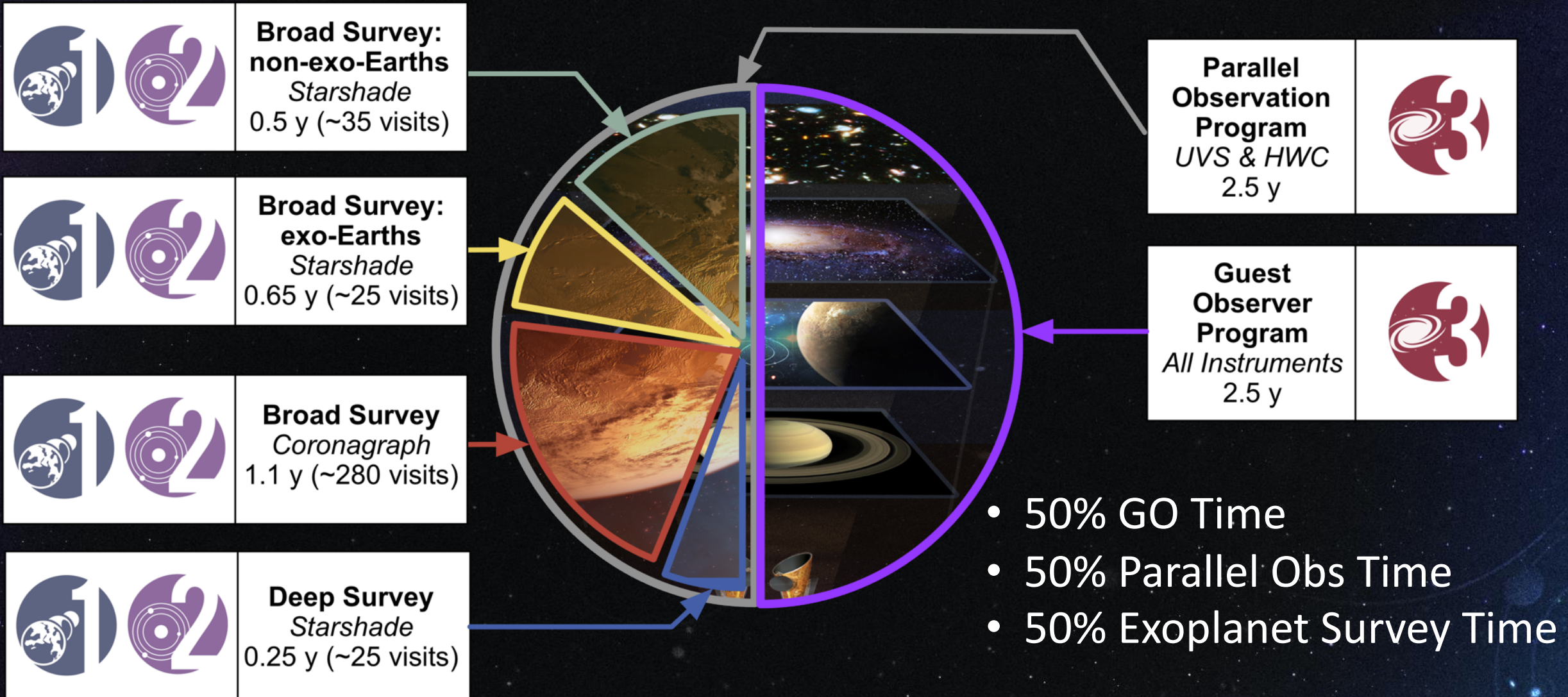
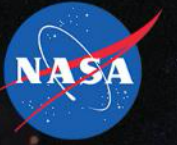
Seek out nearby worlds and
explore their habitability



Map out nearby planetary systems
and understand their diversity.



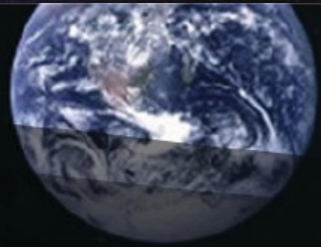
Enable new explorations of
systems in the UV to near-IR



HabEx



Baseline Architecture



Inner working
angle (IWA)

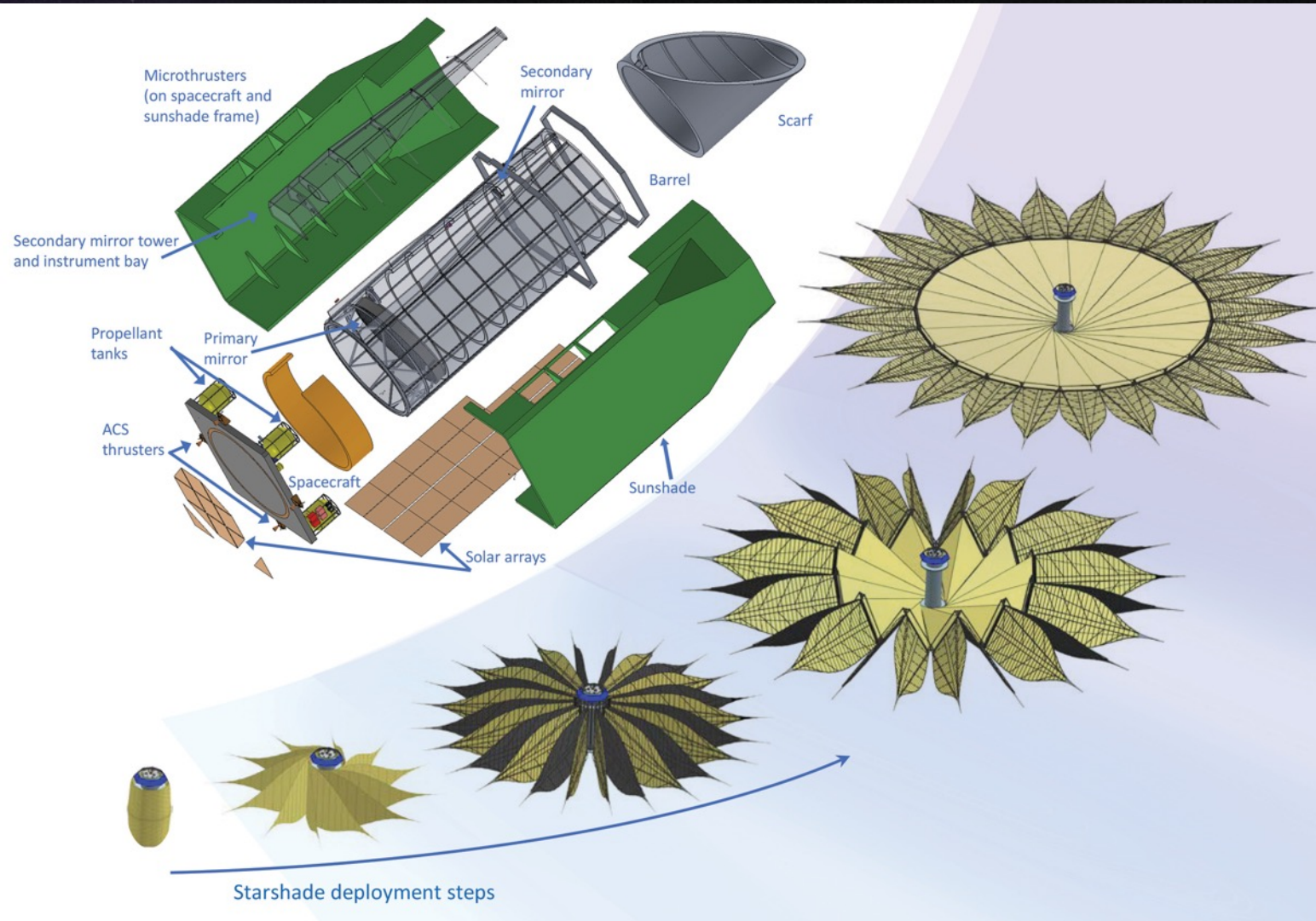
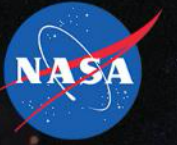
76,600 km separation

Telescope aperture
diameter 4 m

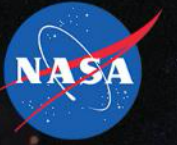


Starshade
diameter 52 m

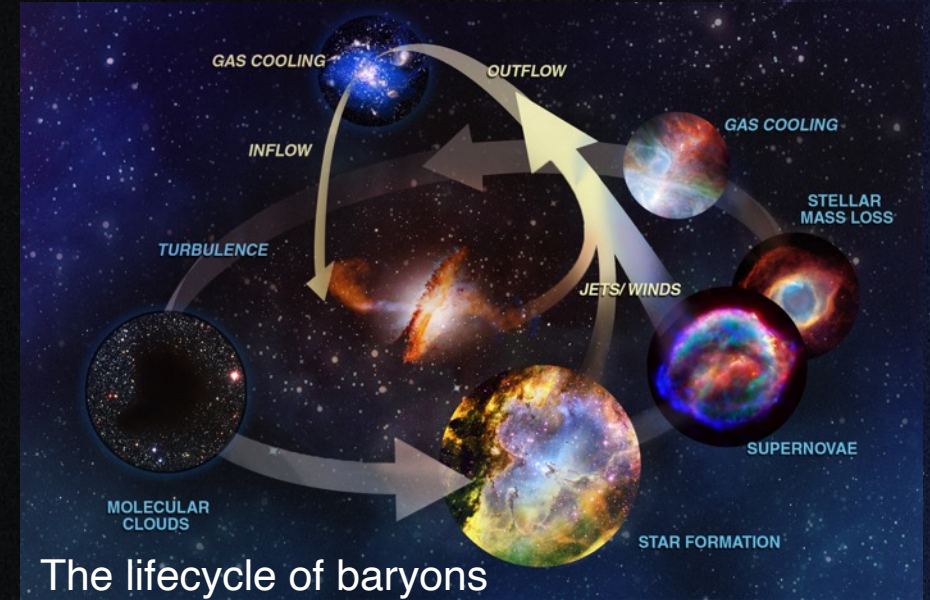
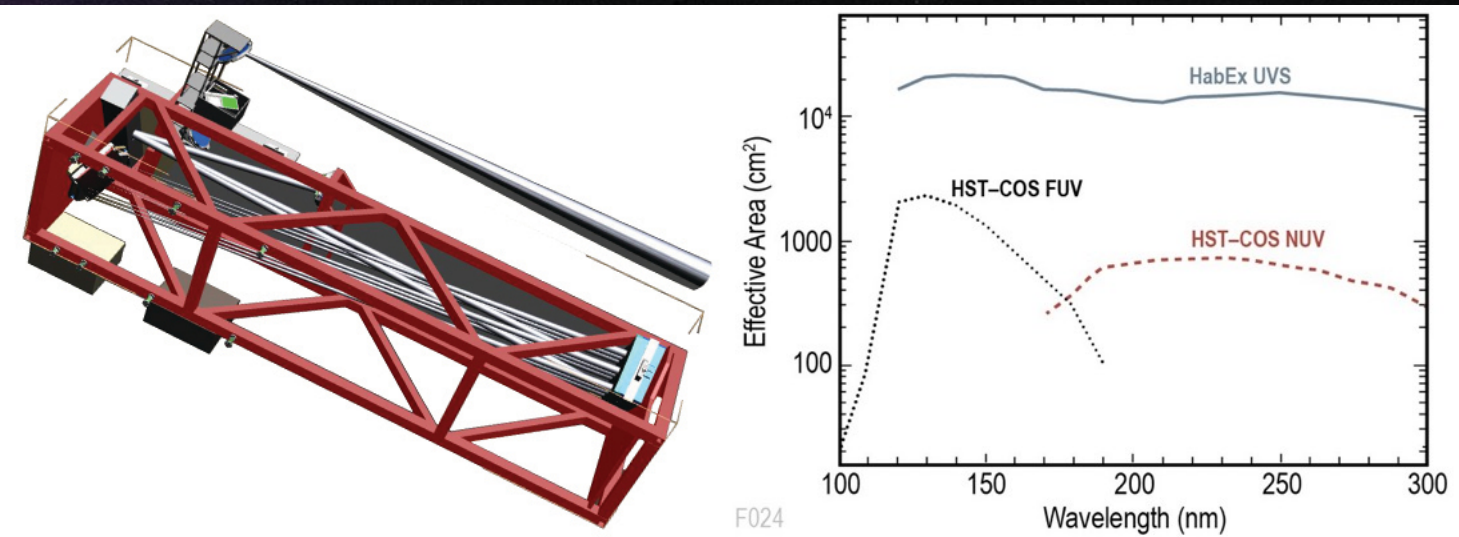




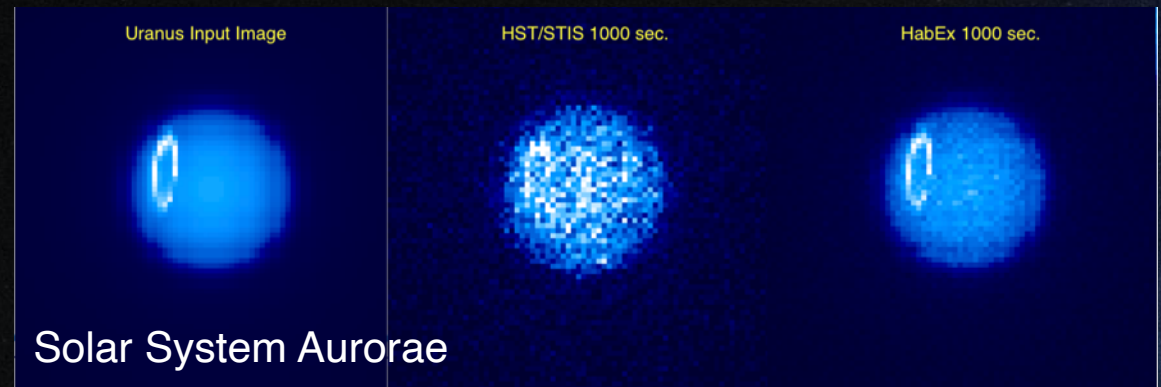
	Habitable Exoplanet Observatory (HabEx)
Mission Duration	5 years (10 years consumables)
Orbit	Earth-Sun L2 Halo
Telescope Aperture	4-meter unobscured
Telescope Type	Off-axis three-mirror anastigmat
Primary Mirror	4-meter monolith; glass-ceramic substrate; Al+MgF2 coating
Instruments (4)	Exoplanet science: Coronagraph, Starshade Observatory science: UV Spectrograph, Workhorse Camera
Attitude Control	Slewing: hydrazine thrusters; Pointing: microthrusters



High-resolution UV imaging and spectroscopy

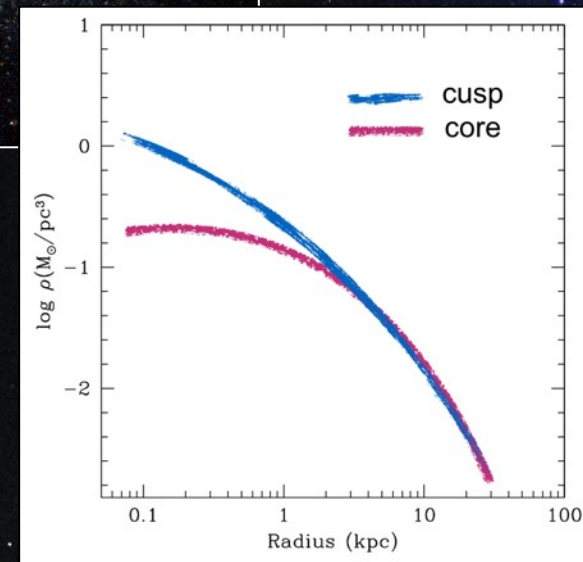
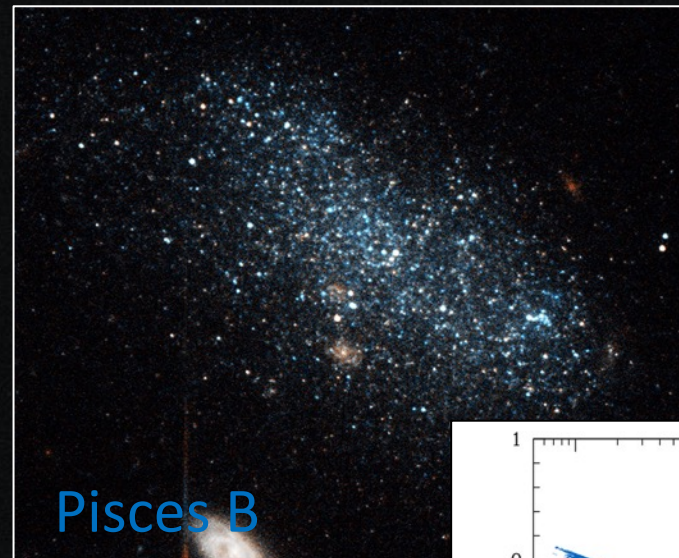
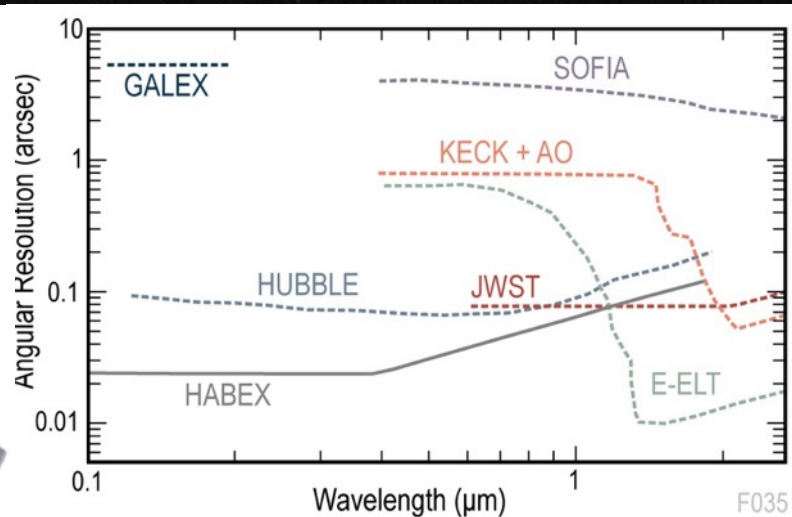
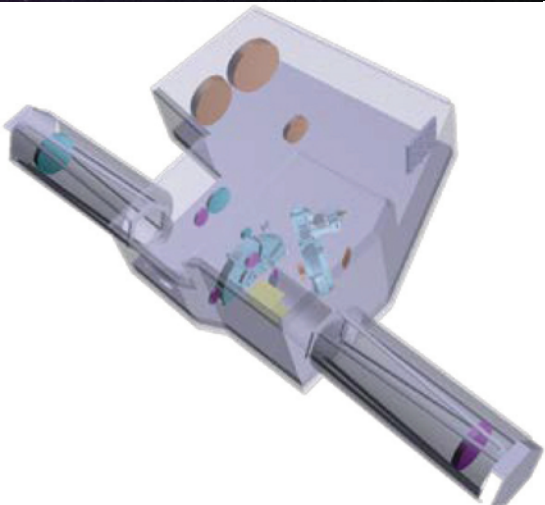


Field-of-view	3 x 3 arcmin ²
Wavelength bands	20 bands covering 0.115 to 0.32 μ m
Spectral resolutions	60,000; 25,000; 12,000; 6,000; 3,000; 1,000; 500
Telescope resolution	Diffraction limited at 0.4 μ m
Detector	3x5 MCP array, 100mm sq each





Multi-purpose optical to near-IR wide-field imaging and spectroscopy



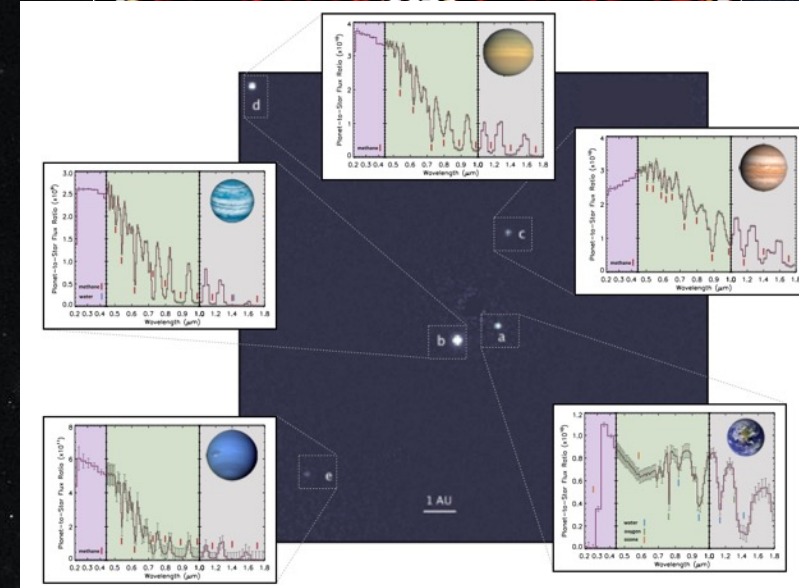
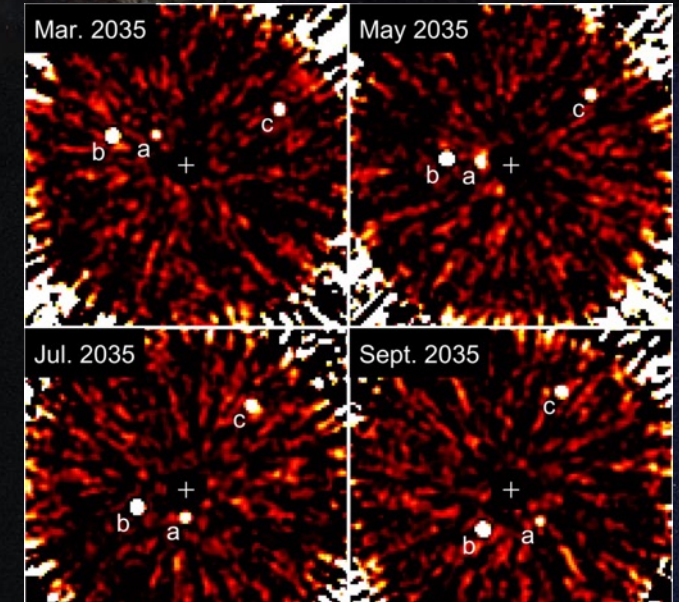
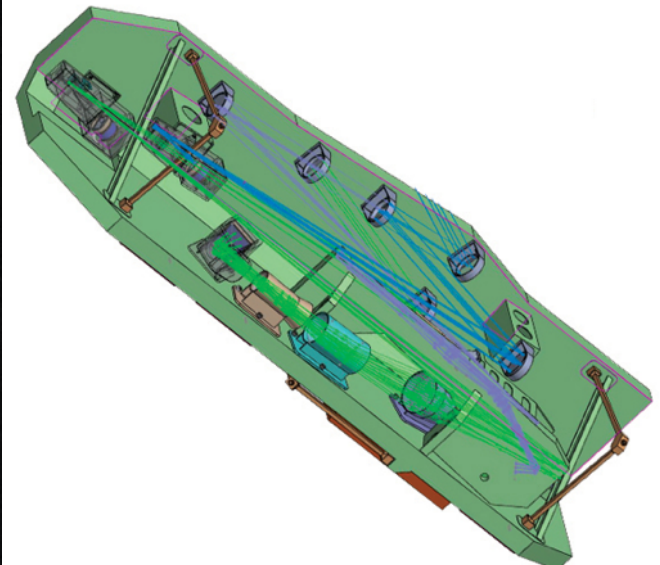
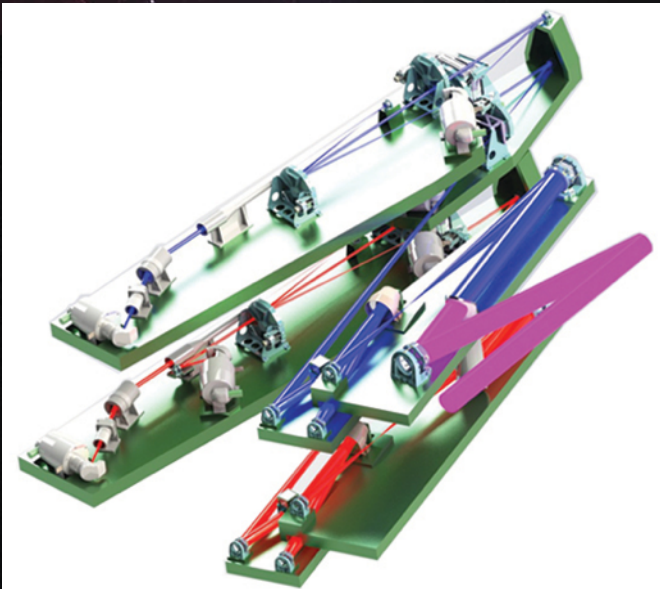
	Vis Channel	near-IR Channel
Field-of-view	3 x 3 arcmin ²	3 x 3 arcmin ²
Wavelength bands	0.37 – 0.975 μm	0.975 – 1.8 μm
Spectral resolution	1000	1000
Telescope resolution	30.9 mas	49 mas
Detectors	3x3 CCD203	2x2 H4RG10

Dark matter in dwarf galaxies



HabEx Coronagraph	Blue Channel	Red Channel	Near-IR Channel
Field-of-view	1.5x1.5 arcsec ²	2.2x2.2 arcsec ²	3.1x3.1 arcsec ²
Wavelength bands	0.45 – 0.55 μm 0.55 – 0.67 μm	0.67 – 0.82 μm 0.82 – 1.0 μm	0.975 – 1.8 μm (in 3 20% bands)
Telescope resolution	23 mas @ 0.45 μm	42 mas @ 0.82 μm	49 mas @ 0.95 μm
IWA	56 mas	102 mas	118 mas
OWA	0.74 arcsec	1.11 arcsec	1.57 arcsec
Spectral resolution	140	140	40

HabEx Starshade	UV Channel	Visible Channel	Near-IR Channel
Imaging field-of-view	10 x 10 arcsec ²	12 x 12 arcsec ²	Pupil imager
Wavelength bands	0.2 – 0.45 μm	0.45 – 0.975 μm	0.975 – 1.8 μm
Telescope resolution	21 mas	21 mas	-
IWA _{0.5}	39 mas	58 mas	104 mas
Spectral resolution	7	140	40





9 architectures investigated

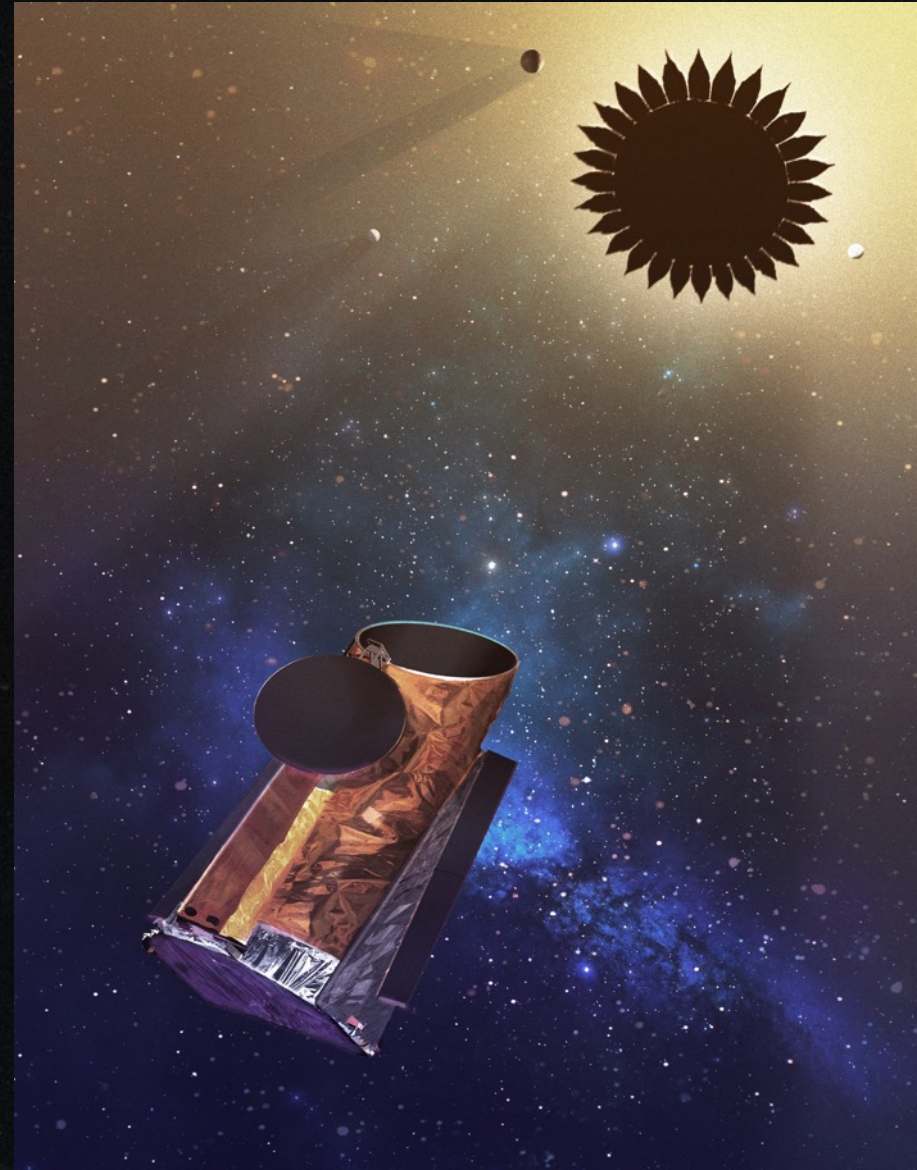
- Three aperture options: 4m; 3.2m; 2.4m.
- All apertures and options include the UVS and HWC instruments.
- Three options for exoplanet instruments per aperture:
 - Hybrid (**H**) with both coronagraph and starshade instruments using an off-axis monolithic primary mirror.
 - Coronagraph instrument only (**C**) using an off-axis monolithic telescope primary mirror.
 - Starshade instrument only (**S**) using an on-axis segmented primary mirror.

		Architecture								
		4.0H	4.0C	4.0S	3.2H	3.2C	3.2S	2.4H	2.4C	2.4S
Science Objective	O1. Exo-Earth candidates around nearby sunlike stars?									
	O2. Water vapor in rocky exoplanet atmospheres?									
	O3. Biosignatures in rocky exoplanet atmospheres?									
	O4. Surface liquid water on rocky exoplanets?									
	O5. Architectures of nearby planetary systems?									
	O6. Exoplanet atmospheric variations in nearby planetary systems?									
	O7. Water transport mechanisms in nearby planetary systems?									
	O8. Debris disk architectures in nearby planetary systems?									
	O9. Lifecycle of baryons?									
	O10. Origins of the elements?									
	O11. Discrepancies in measurements of the cosmic expansion rate?									
	O12. The nature of dark matter?									
	O13. Formation and evolution of globular clusters?									
	O14. Potentially habitable conditions on rocky planets around M-dwarfs?									
	O15. Mechanisms responsible for transition disk architectures?									
	O16. Physics driving star-planet interactions, e.g. auroral activity?									

Not yet finalized



- What would YOU do with HabEx?!
- The HabEx Team is soliciting 2-page community white papers to include in the HabEx Final Report.
- Email Rashied Amini for a template:
Rashied.Amini@jpl.nasa.gov



HabEx





New Worlds, New Horizons

in Astronomy and Astrophysics

NATIONAL RESEARCH COUNCIL
OF THE NATIONAL ACADEMIES

- The National Research Council convenes experts to define the ground- and space-based Astrophysics priorities for the next 10 years.
- For large space-based telescopes –
 - The 2000 Decadal Survey prioritized the James Webb Space Telescope (JWST), which will launch in ~2021.
 - The 2010 Decadal Survey prioritized the Wide Field Infrared Survey Telescope (WFIRST), which will launch in ~2025.
 - The Habitable Exoplanet Observatory (**HabEx**) is one of four “flagship” mission studies currently being undertaken by NASA for consideration in the **2020 Decadal Survey**. If prioritized, HabEx would launch in the mid-to-late 2030’s.
- The other studies are:
 - The Origins Space Telescope (OST)
 - The Large UV, Optical, IR Telescope (LUVOIR)
 - Lynx



“Develop an optimal* mission concept for characterizing our nearest planetary systems, and detecting and characterizing a handful of ExoEarths.”

“Given this optimal* concept, maximize the astrophysics science potential without sacrificing the primary exoplanet science goals.”

What does optimal mean?

- Maximizing the science yield while maintaining feasibility, i.e., adhering to expected constraints.
- Constraints include: Cost, technology (risk), time to develop mission.